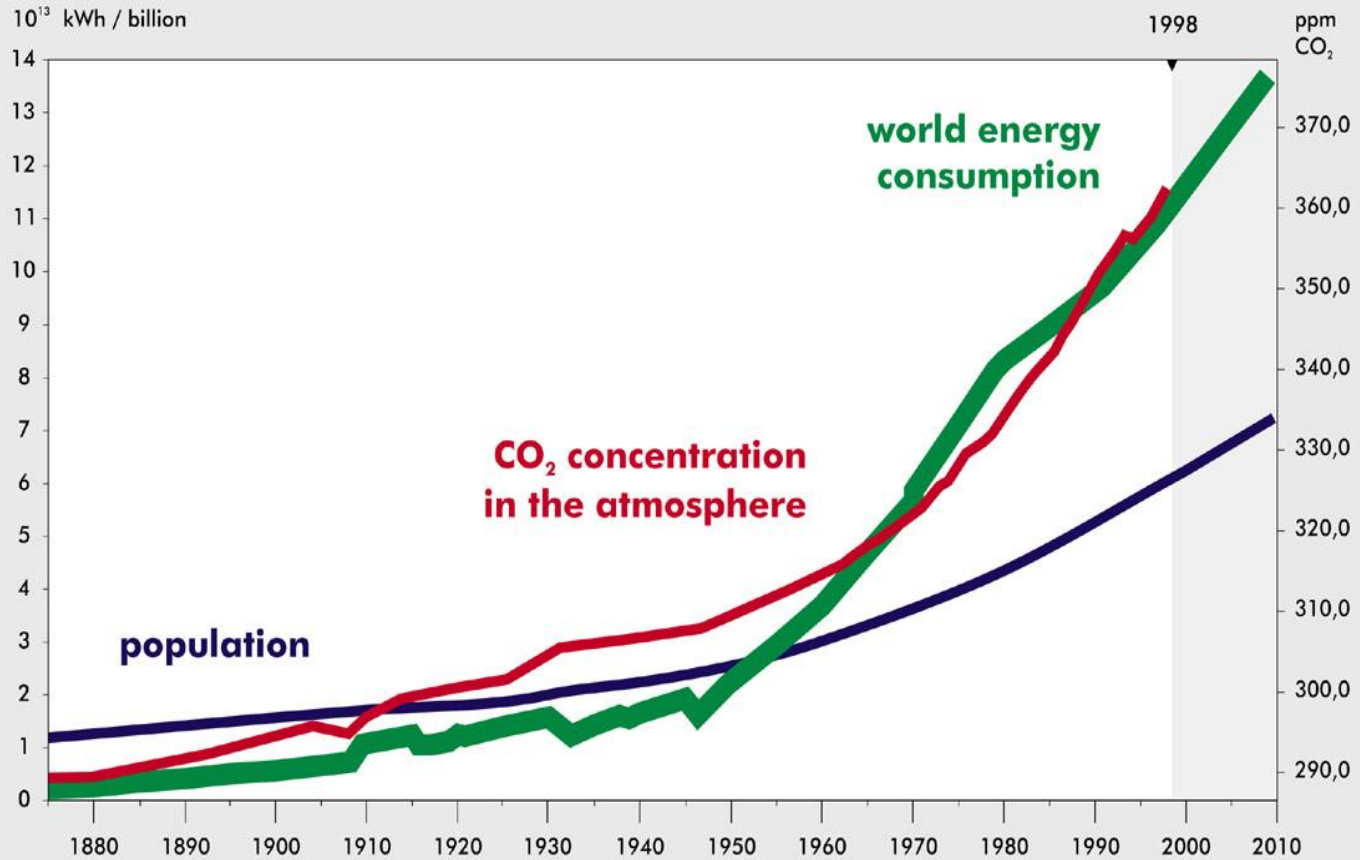


Solar Energy Exploitation Systems



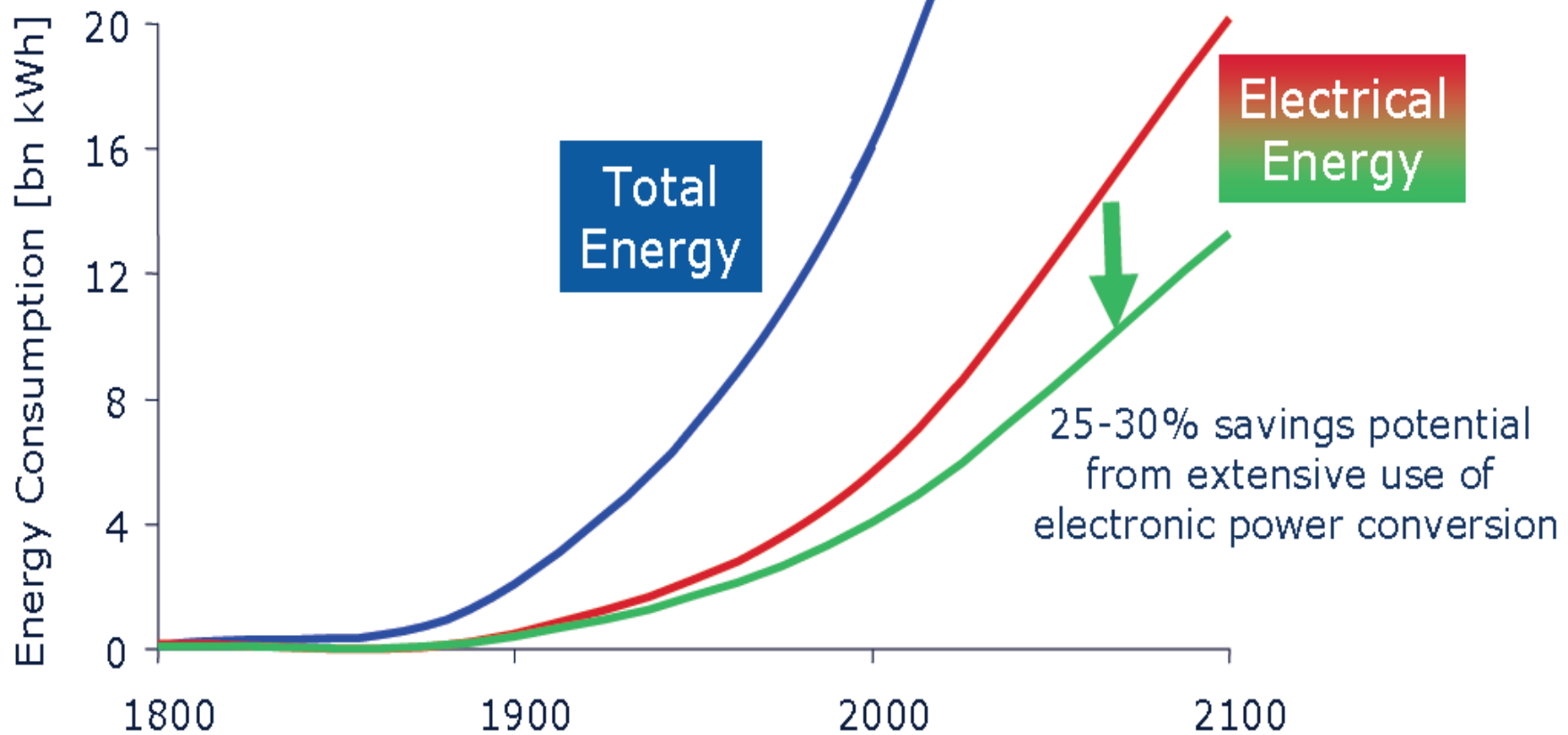
Course structure - lectures

1. Solar energy and basic forms of its exploitation
2. Photovoltaic phenomena
3. Solar cells and their characteristics
4. Monocrystalline, polycrystalline and thin film solar cells
5. Solar cells from other semiconductor materials
6. Construction and fabrication of solar cells
7. Construction and fabrication of modules
8. Basic types of photovoltaic systems
9. Energy storage for photovoltaic systems
10. Stand alone and on-grid systems
11. Applications of Photovoltaic systems
12. Basic economic and ecological aspects
13. Solar thermal systems
14. Conversion solar energy to electrical energy



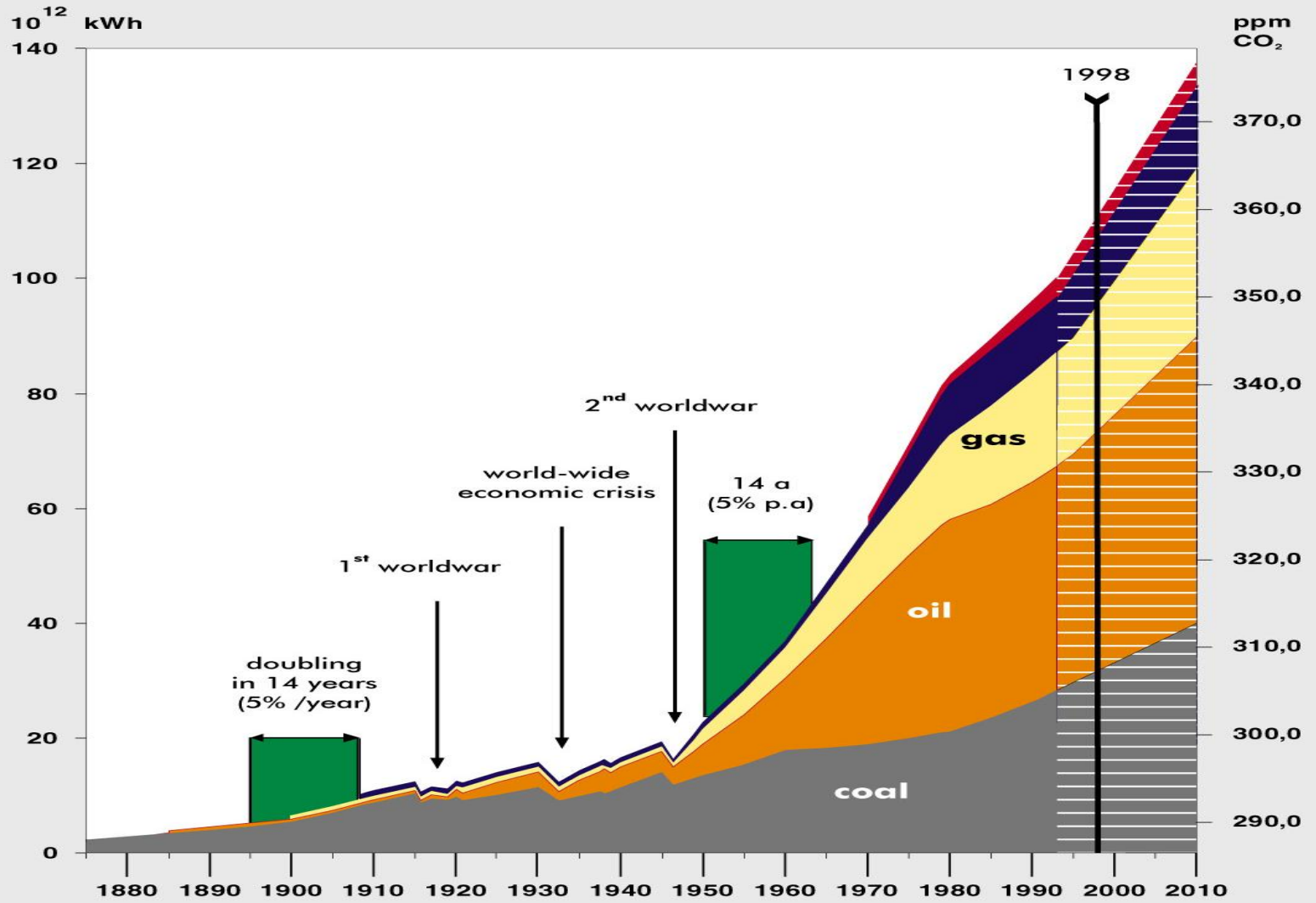
Growth rates for population, energy consumption and CO₂ concentration in the atmosphere

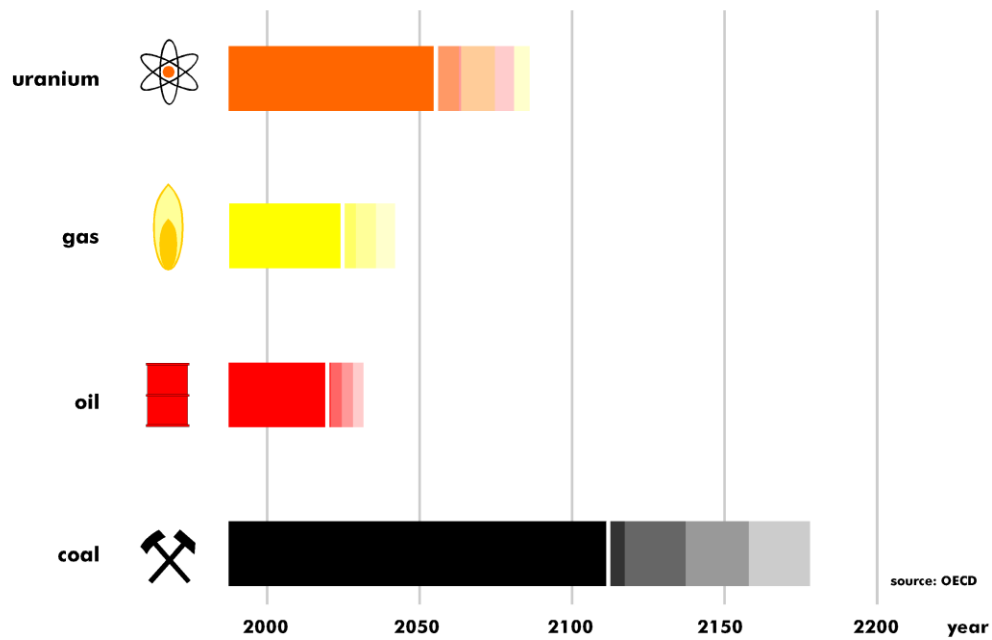
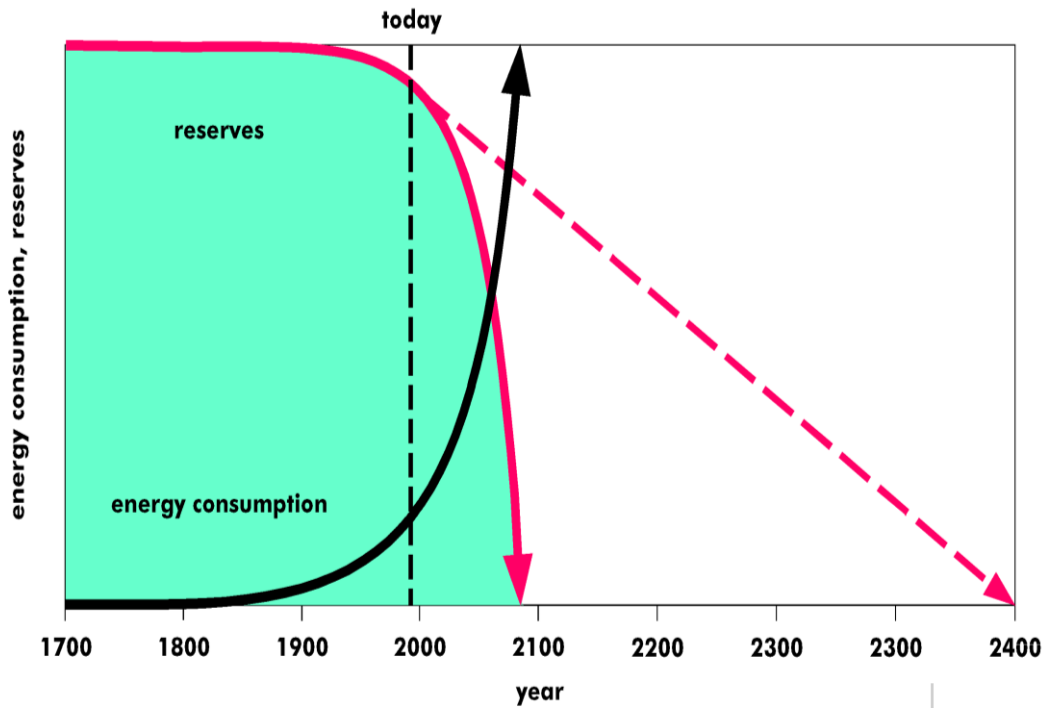




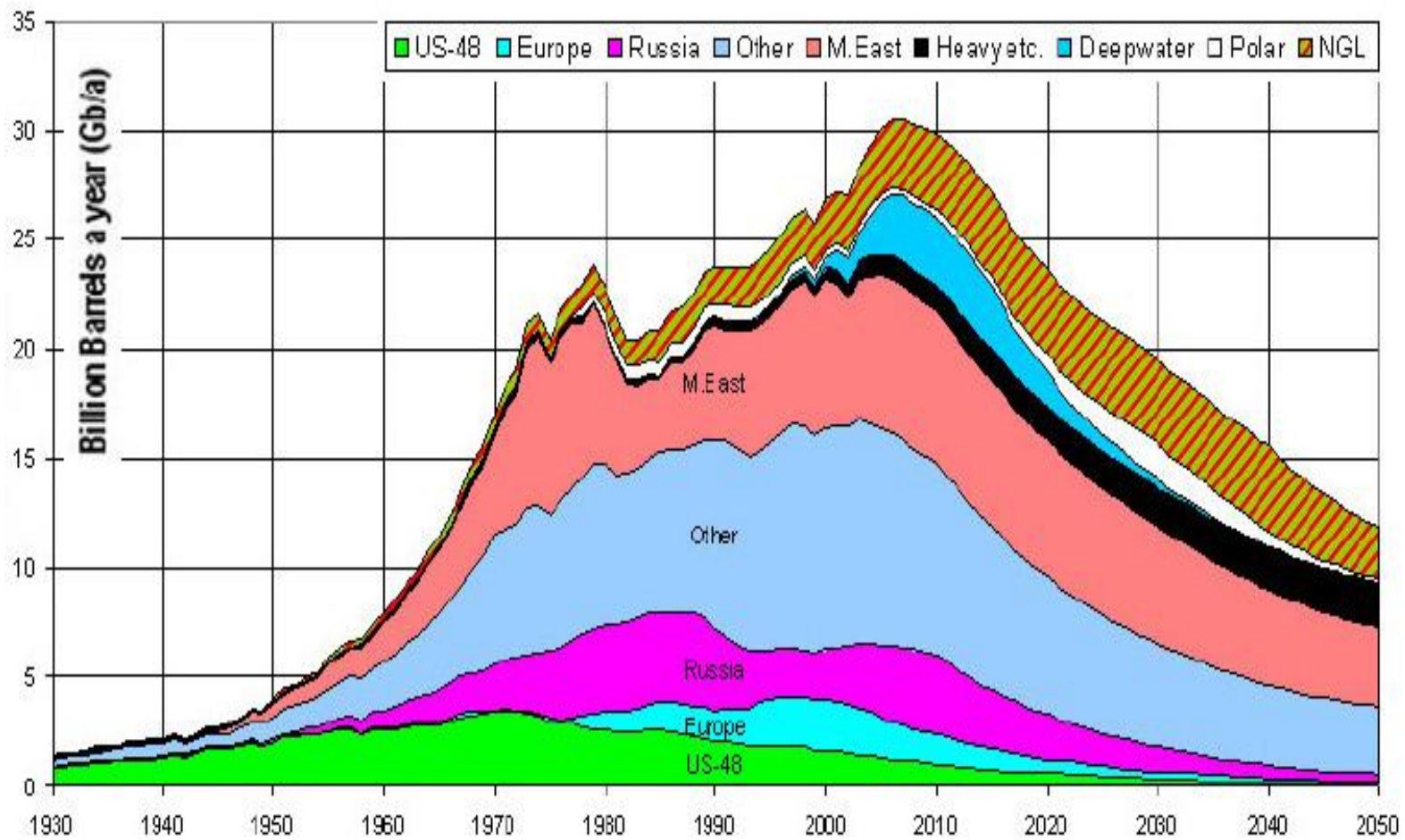
Energy Demand increases at 2.6 % annually

source: Interatom/Shell 3/1992

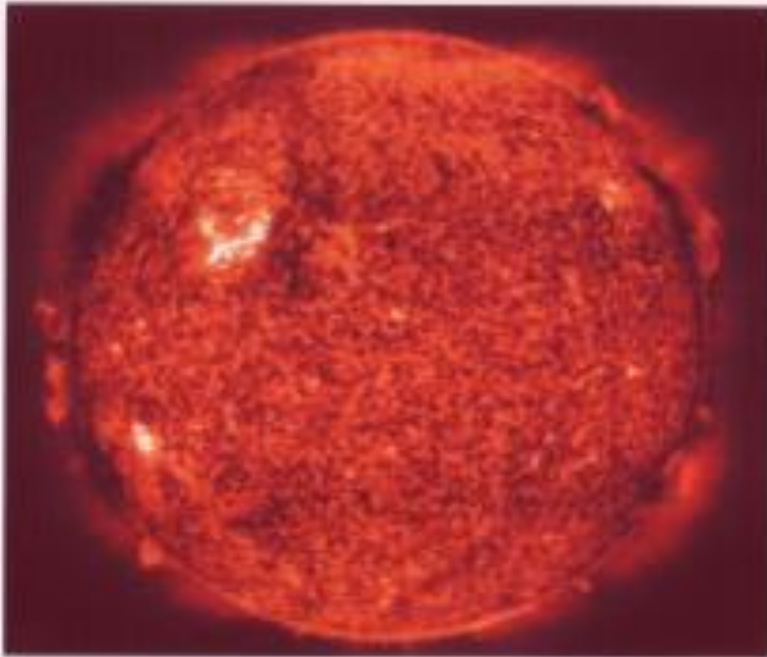




Scarcity of different energy sources



The Sun



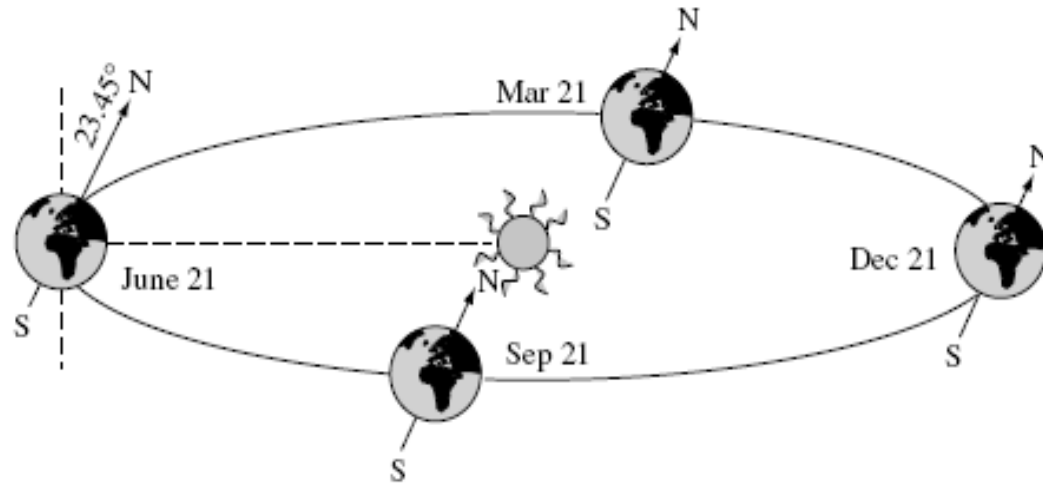
$$r_{\text{S}} = 0.695 \times 10^9 \text{ m}$$

Surface temperature $\sim 6000 \text{ K}$

Distance from the Earth

$$R = 1.496 \times 10^{11} \text{ m } (\pm 1.7\%)$$

Energy of Solar Radiation

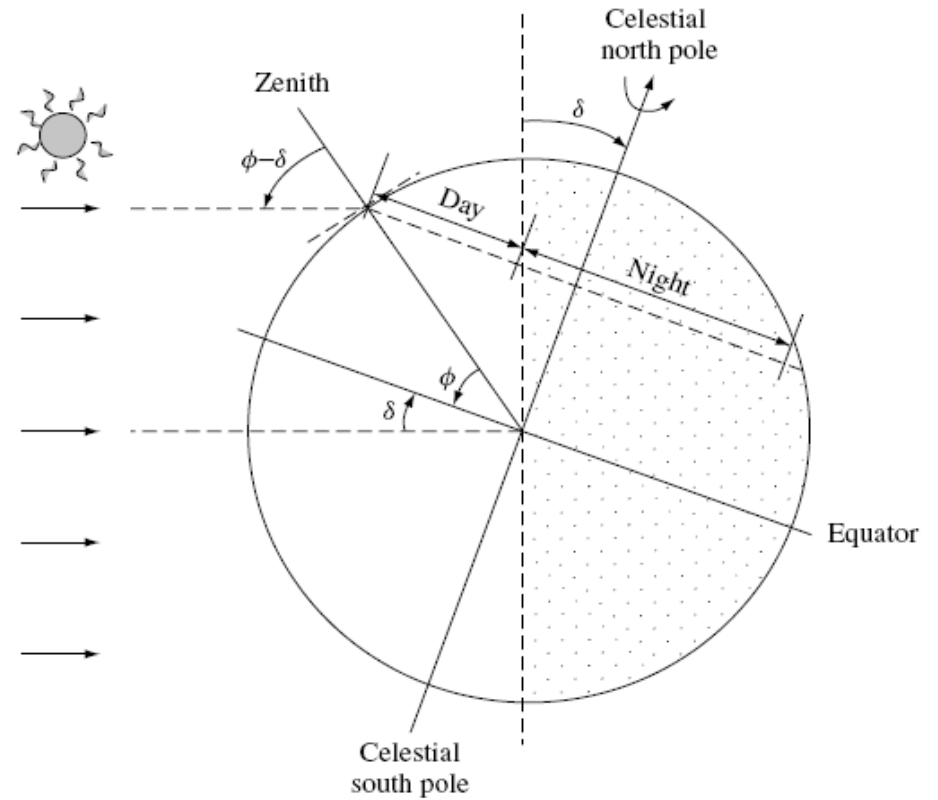
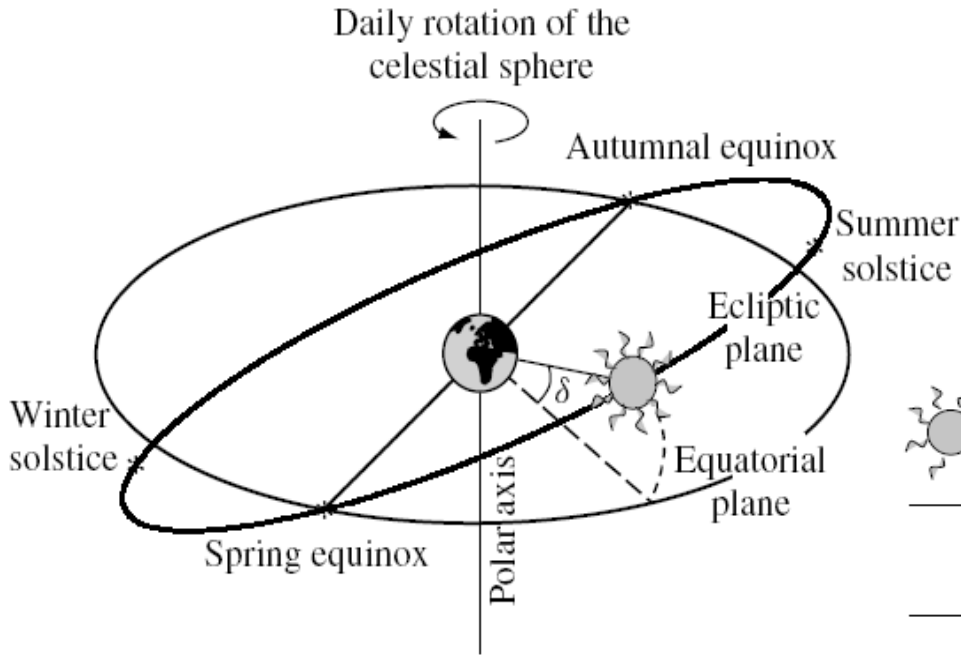


$$r = r_0 \left[1 + 0.017 \sin \left(\frac{360(d_n - 93)}{365} \right) \right]$$

$$r_0 = 1.496 \times 10^8 \text{ km}$$

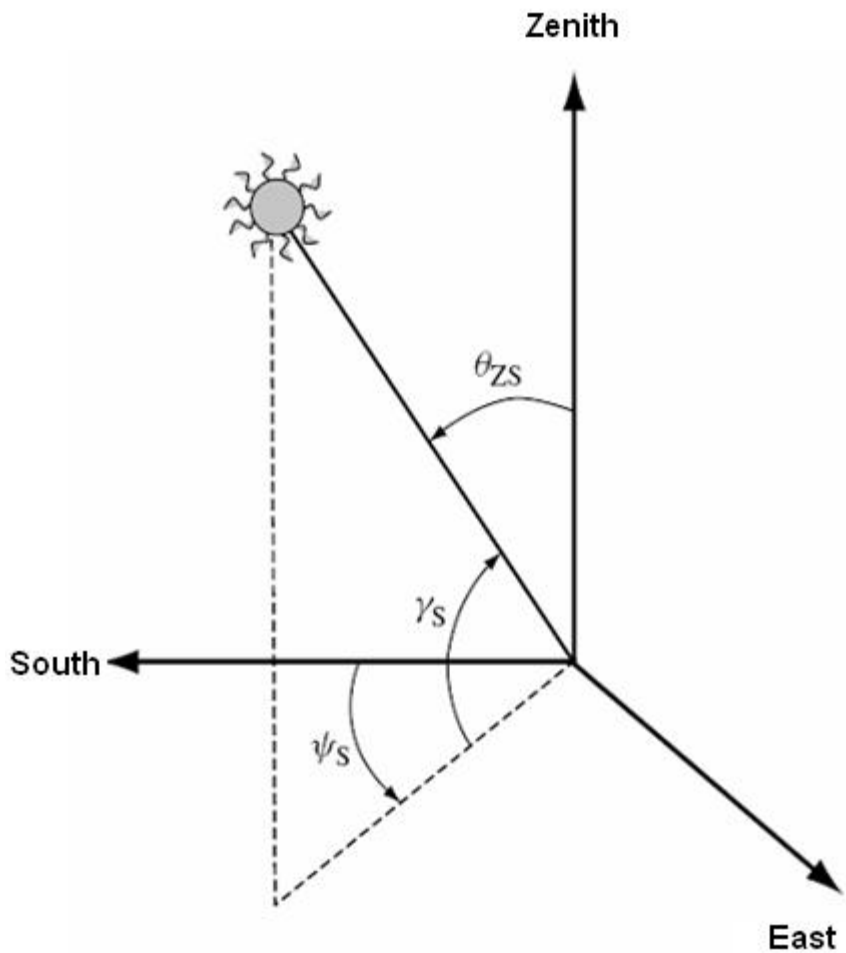
excentrita

$$\varepsilon_0 = (r_0/r)^2 = 1 + 0.033 \cos \left(\frac{360d_n}{365} \right)$$



solar declination δ .

$$\delta = 23.45^\circ \sin \left[\frac{360(d_n + 284)}{365} \right]$$



zenith angle, θ_{ZS}

solar azimuth, ψ_S ,

solar altitude, γ_S

geographic latitude, Φ

$$\cos \theta_{ZS} = \sin \delta \sin \phi + \cos \delta \cos \phi \cos \omega = \sin \gamma_S$$

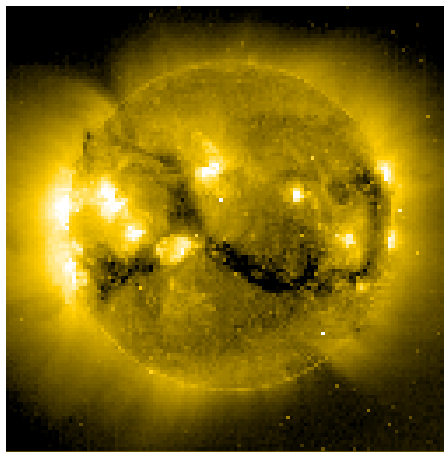
$$\cos \psi_S = \frac{(\sin \gamma_S \sin \phi - \sin \delta)}{\cos \gamma_S \cos \phi} [\text{sign}(\phi)]$$

sunrise angle, ω_S ,

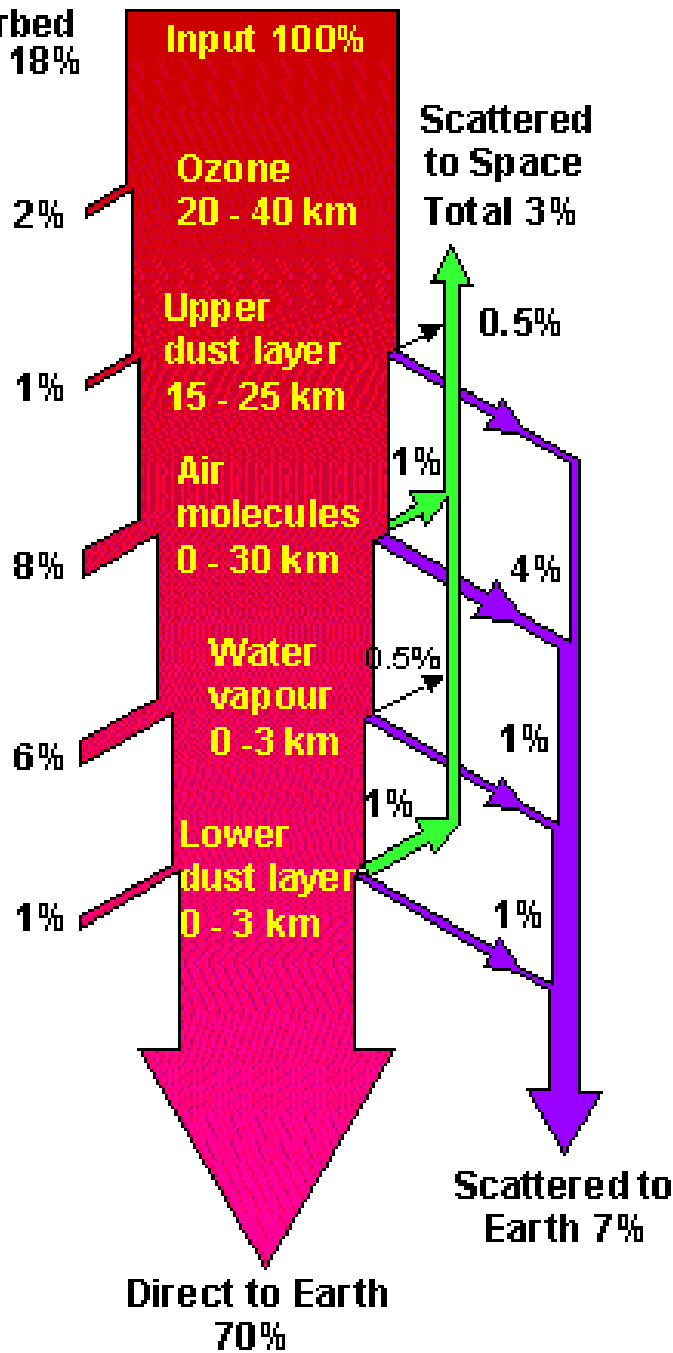
$$\omega_S = -\arccos(-\tan \delta \tan \phi)$$

Declination and extraterrestrial irradiation values for the characteristic day of each month

Month	Date	d_n	δ [degrees]	$B_{0d}(0) = B_{0dm}(0)$, in [Wh/m ²]			
				$\phi = 30^\circ$	$\phi = 60^\circ$	$\phi = -30^\circ$	$\phi = -60^\circ$
January	17	17	-20.92	5 907	949	11 949	11 413
February	14	45	-13.62	7 108	2 235	11 062	9 083
March	15	74	-2.82	8 717	4 579	9 531	5 990
April	15	105	+9.41	10 225	7 630	7 562	3 018
May	15	135	+18.79	11 113	10 171	5 948	1 225
June	10	161	+23.01	11 420	11 371	5 204	605
July	18	199	+21.00	11 224	10 741	5 530	878
August	18	230	+12.78	10 469	8 440	6 921	2 294
September	18	261	+1.01	9 121	5 434	8 835	4 937
October	19	292	-11.05	7 436	2 726	10 612	8 226
November	18	322	-19.82	6 056	1 114	11 754	10 983
December	13	347	-23.24	5 498	613	12 174	12 177

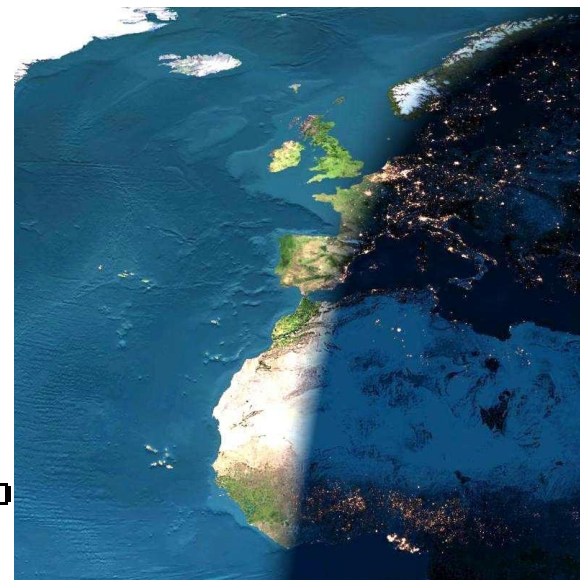


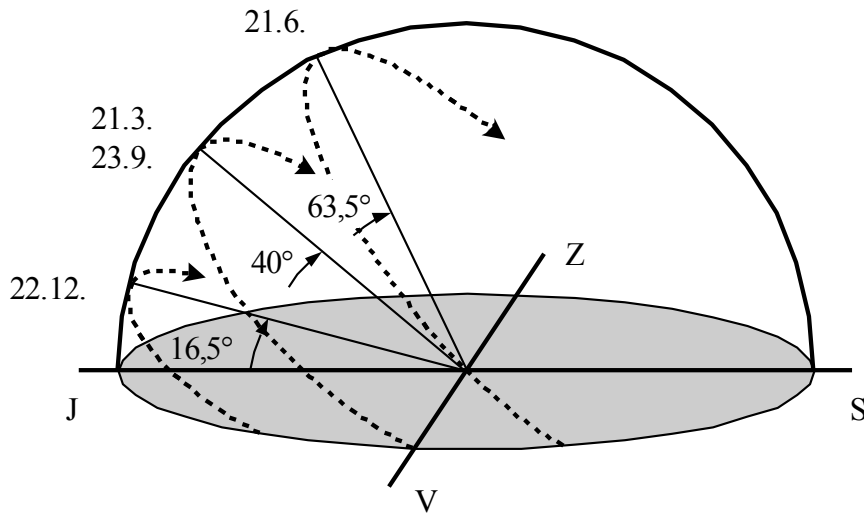
**Absorbed
Total 18%**



1000W/m²

1 – 6MWh/m²

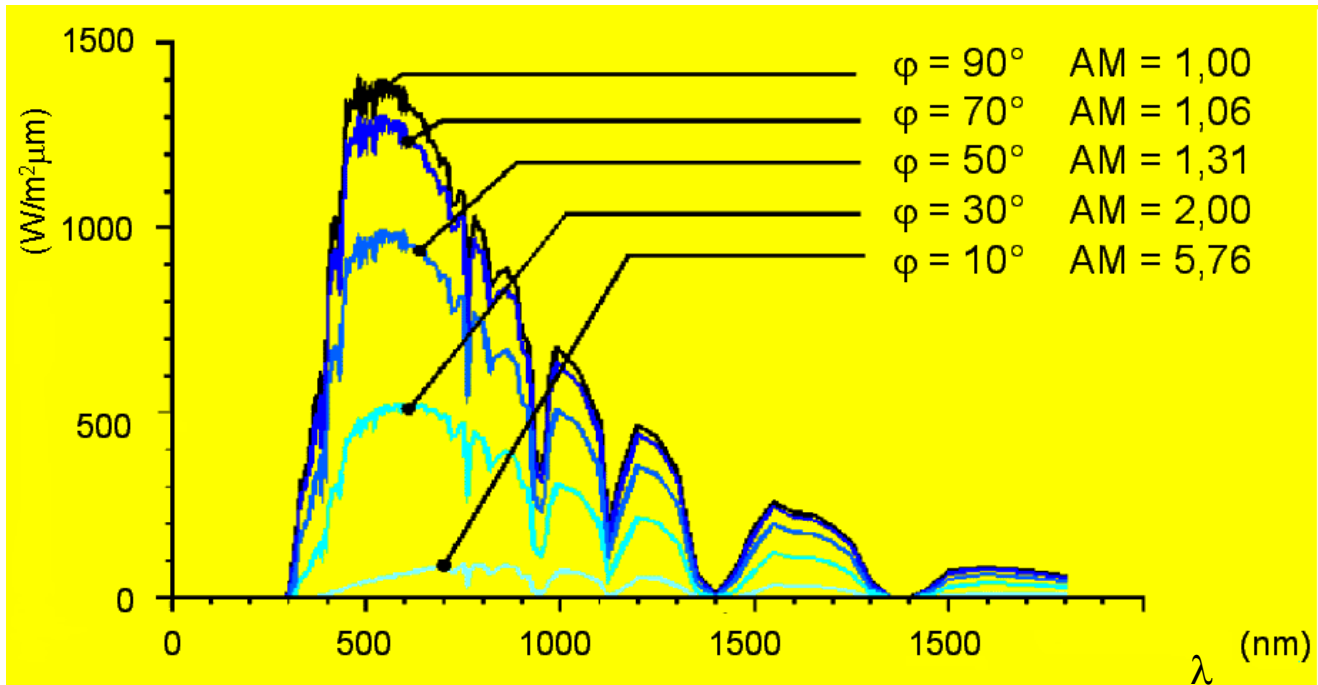


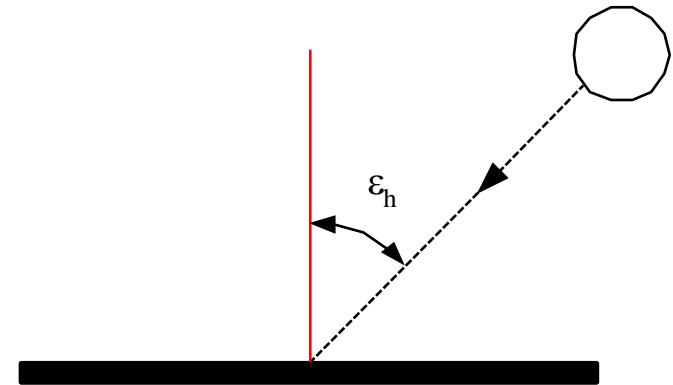
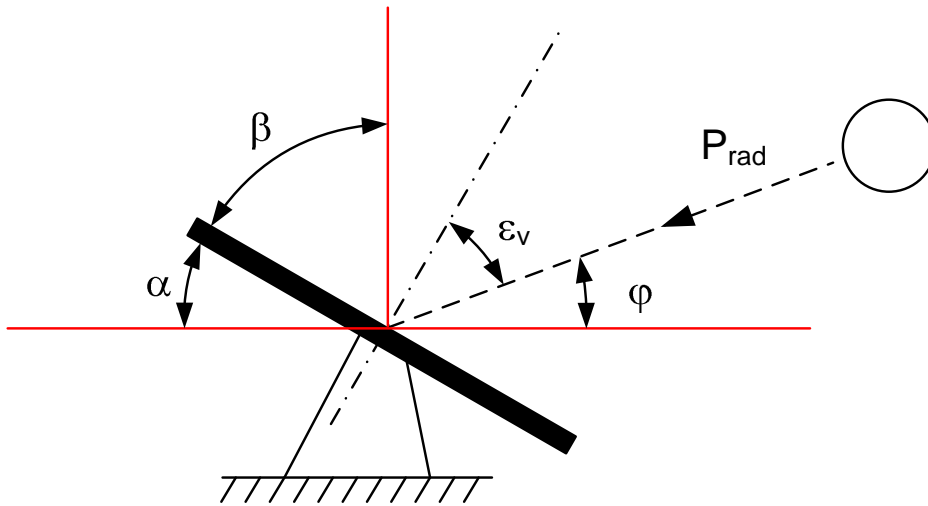
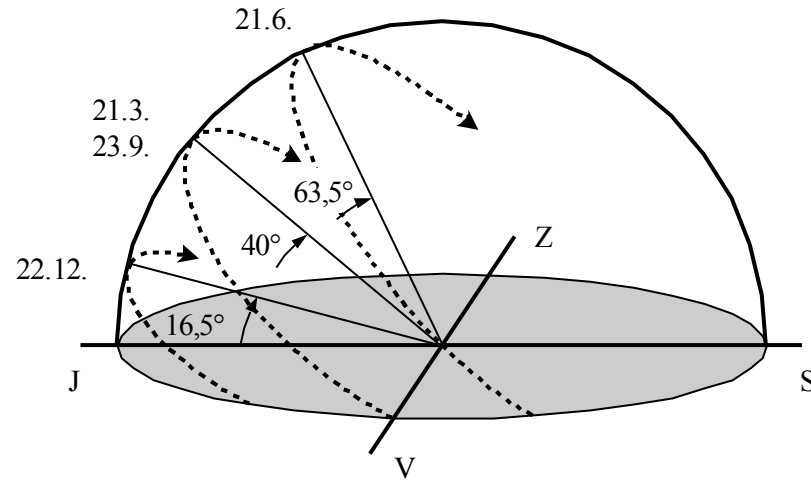


$$B(\varphi) = B_0 (0,7)^{AM}$$

$$AM = \frac{1}{\sin \varphi}$$

Atmospheric Mass coefficient

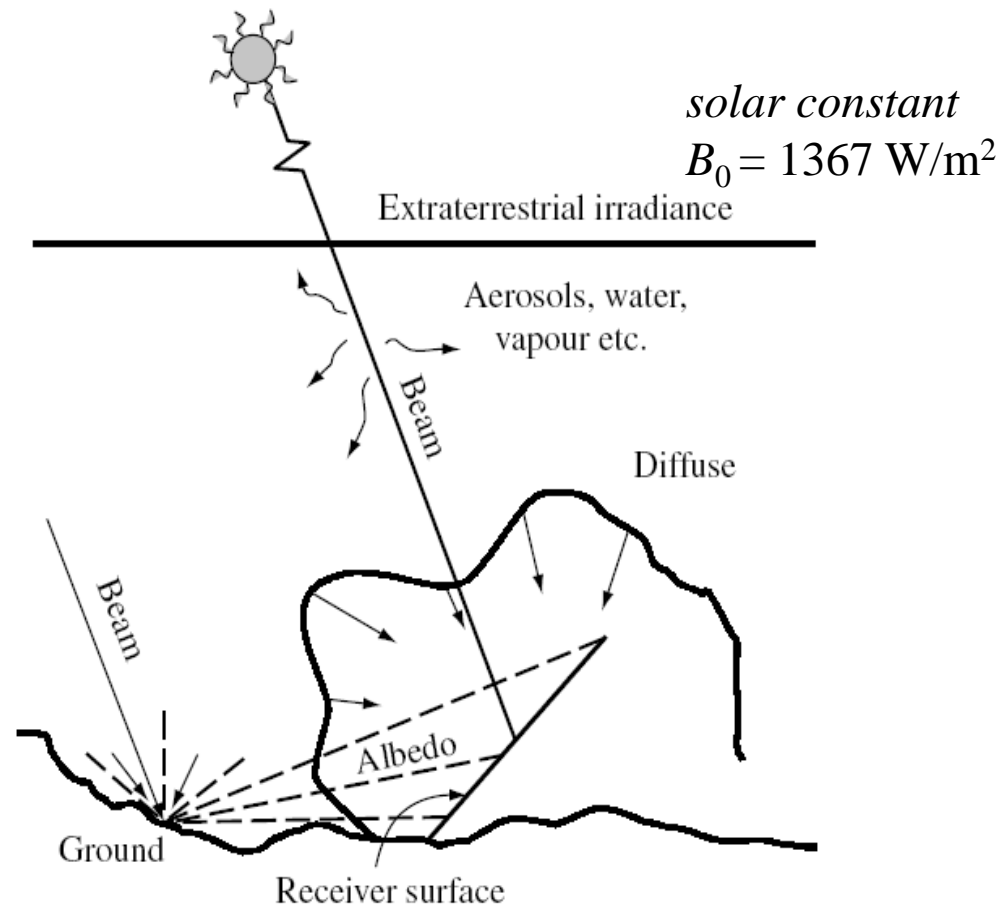




$$P_u(t) = 1367 \times 0.7 \frac{1}{\sin \varphi(t)} \cos \varepsilon_v(t) \cos \varepsilon_h(t)$$

$$W_u = \int_{SR}^{SS} 1367 \times 0.7 \frac{1}{\sin \varphi(t)} \cos \varepsilon_v(t) \cos \varepsilon_h(t) dt$$

Irradiance means density of power falling on a surface, and is measured in W/m^2



beam radiation, made up of beams of light that are not reflected or scattered - ***B***

diffuse radiation, coming from the whole sky apart from the sun's disc - ***D***

albedo radiation is radiation reflected from the ground - ***R***

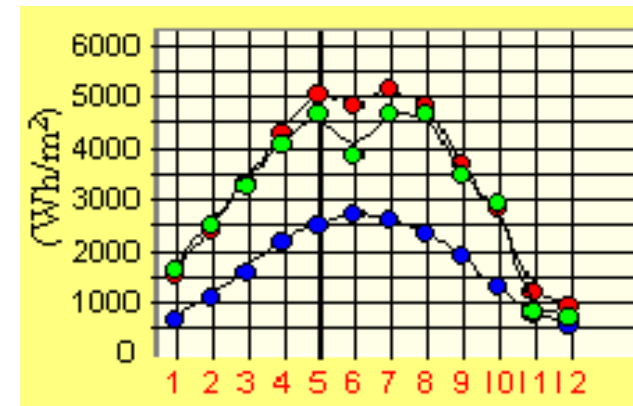
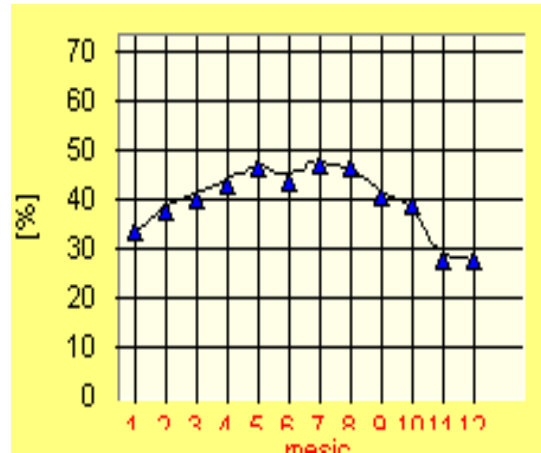
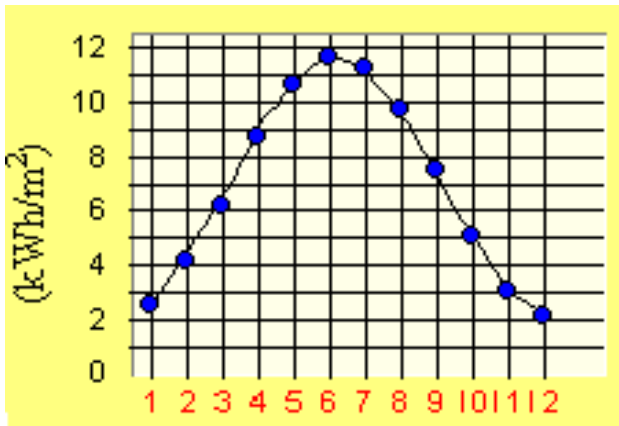
global radiation (direct + diffuse + albedo). **$G = B + D + R$**

There are random variations caused by climatic conditions: cloud cover, dust storms and so on

Monthly mean values of global horizontal daily irradiation, $G_{dm}(0)$

A *clearness index* K_{Tm} , (calculated for each month) gives a measure of the atmospheric transparency.

$$K_{Tm} = \frac{G_{dm}(0)}{B_{0dm}(0)}$$

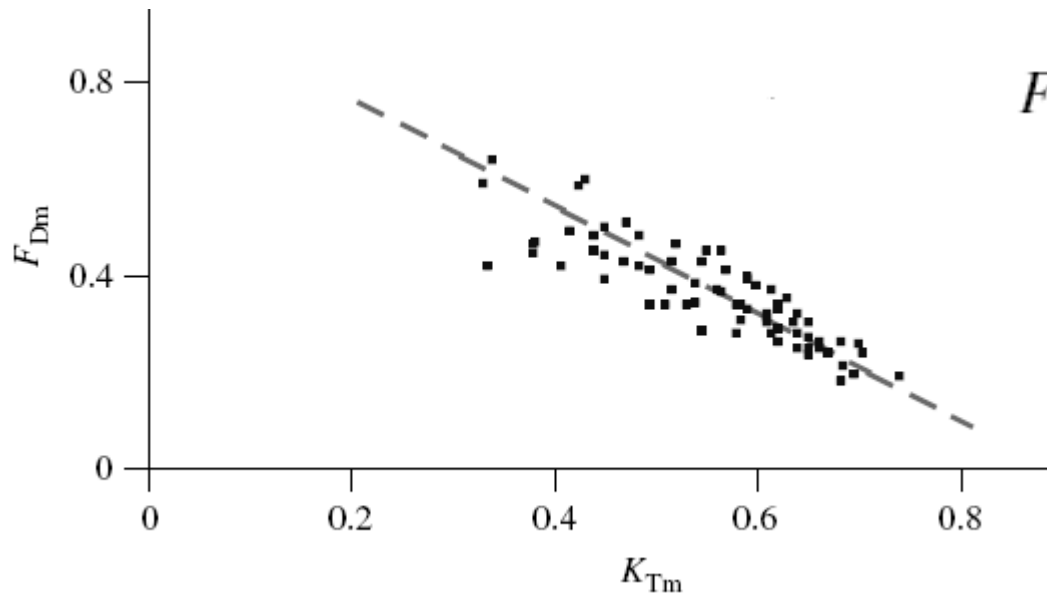


The diffuse fraction of the mean daily global irradiation

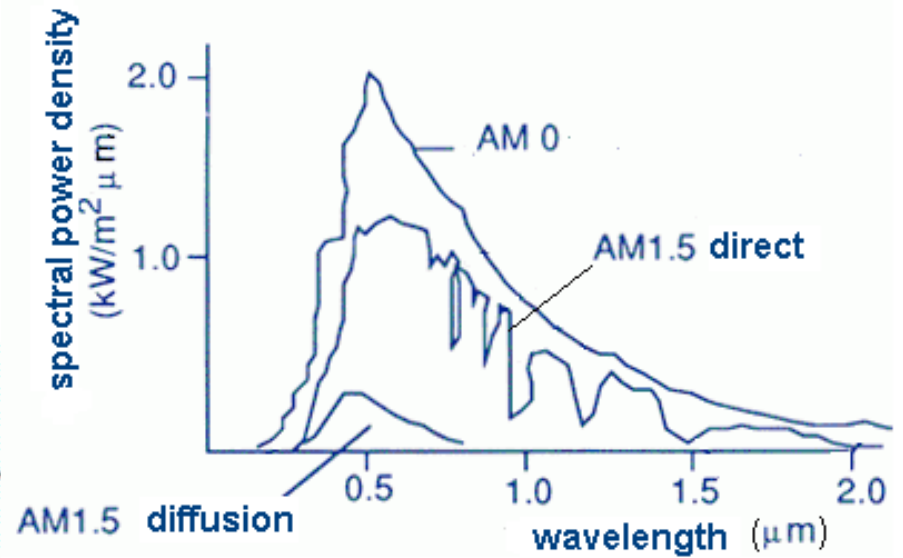
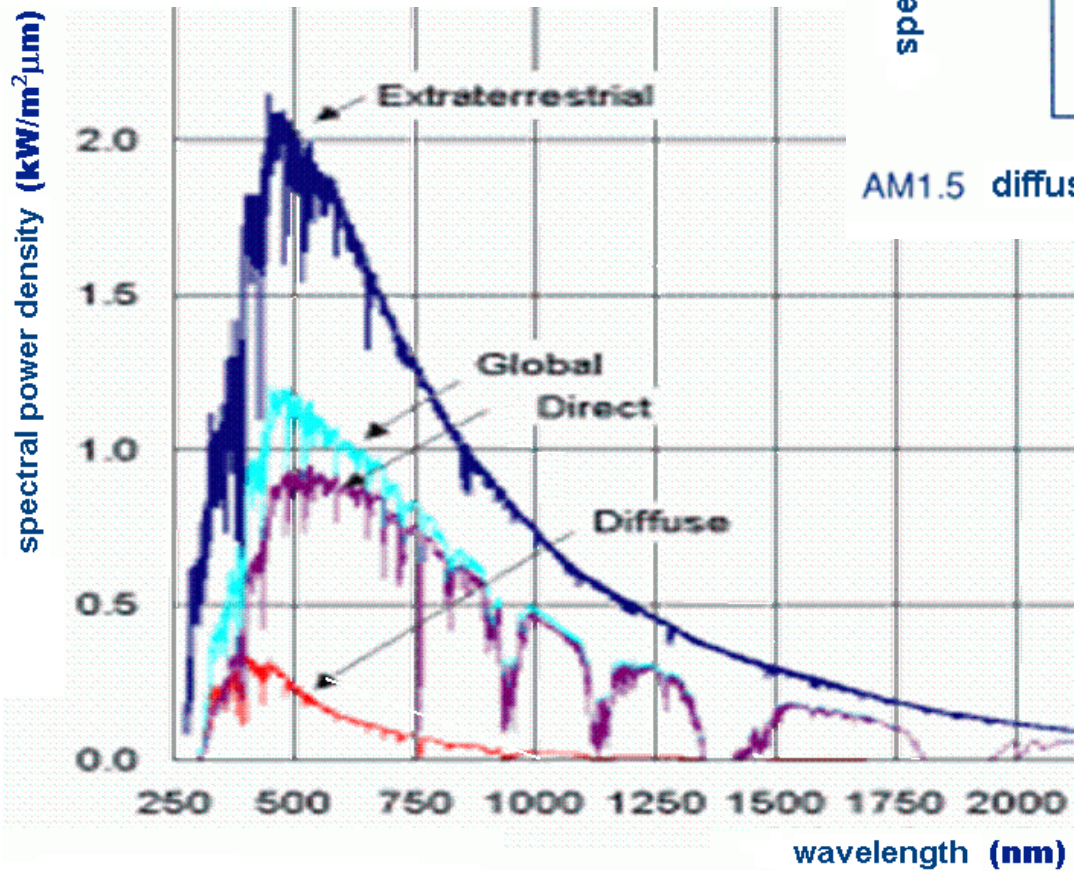
$$F_{Dm} = D_{dm}(0) / G_{dm}(0)$$

The clearness index

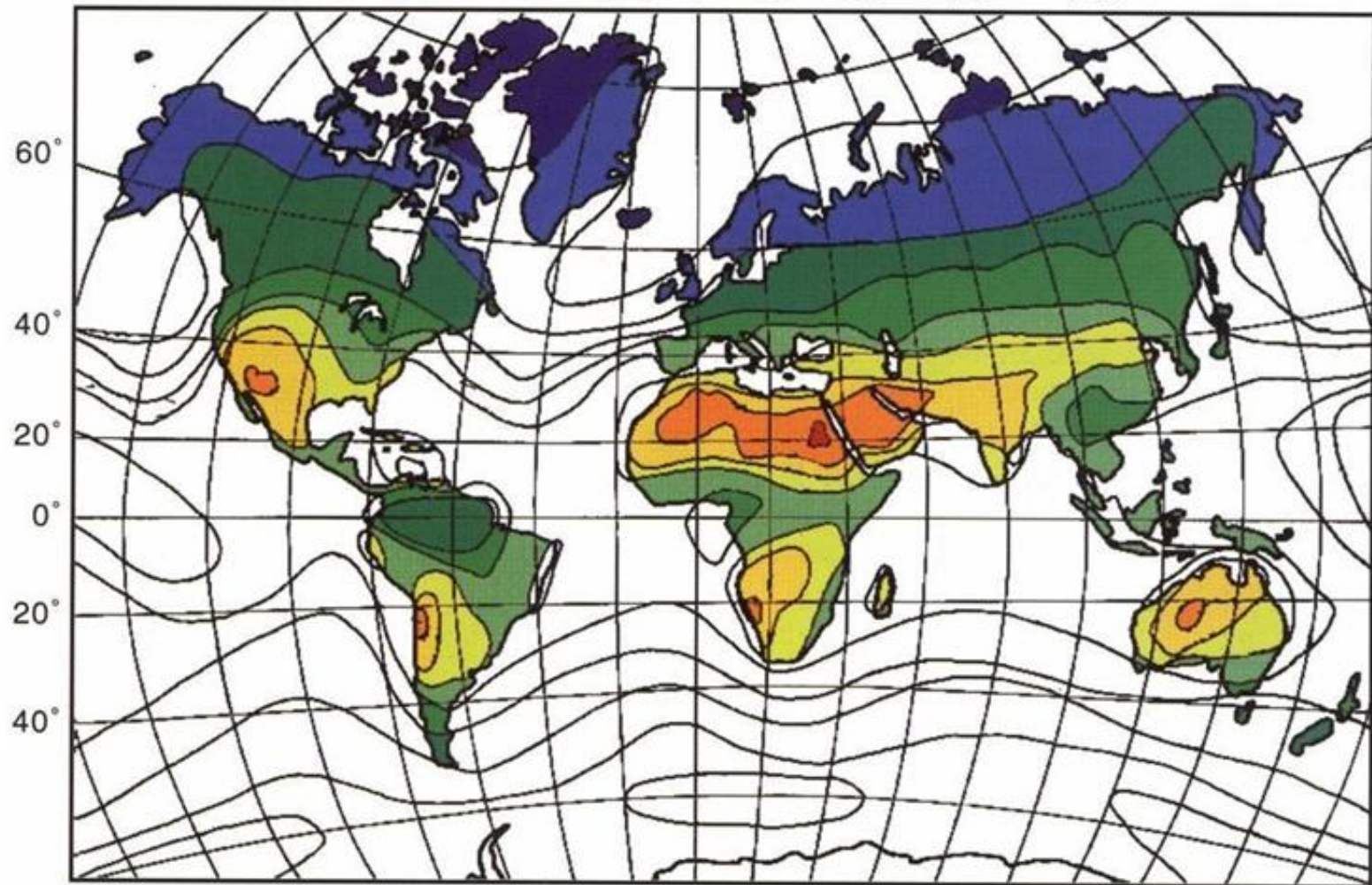
$$K_{Tm} = \frac{G_{dm}(0)}{B_{0dm}(0)}$$



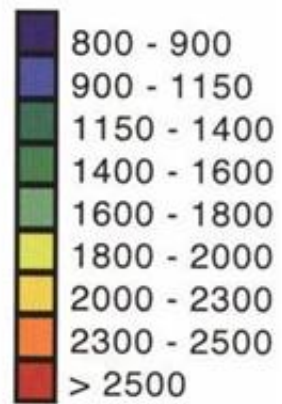
$$F_{Dm} = 1 - 1.13K_{Tm}$$

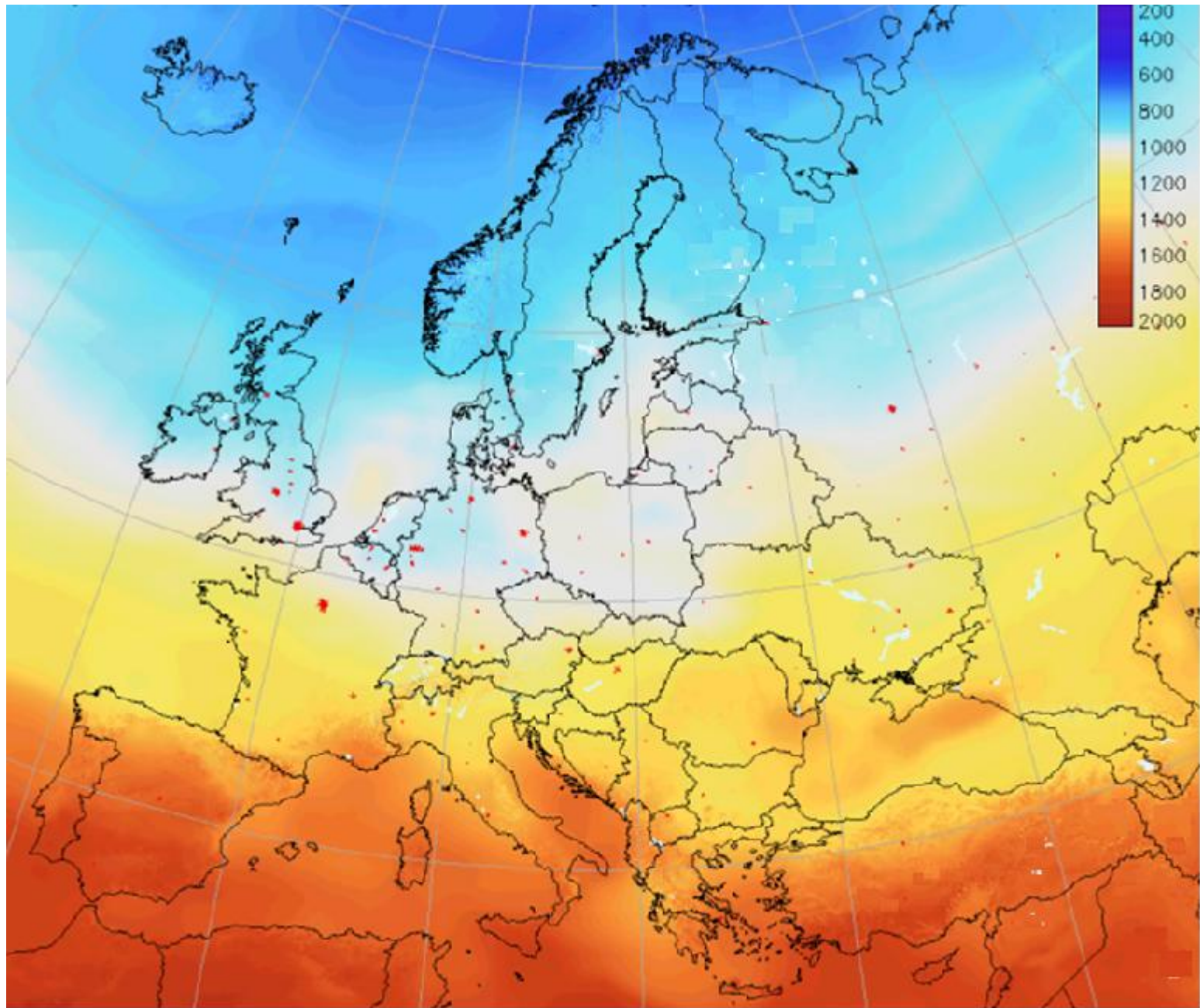


200° 240° 280° 320° 0° 40° 80° 120° 160°

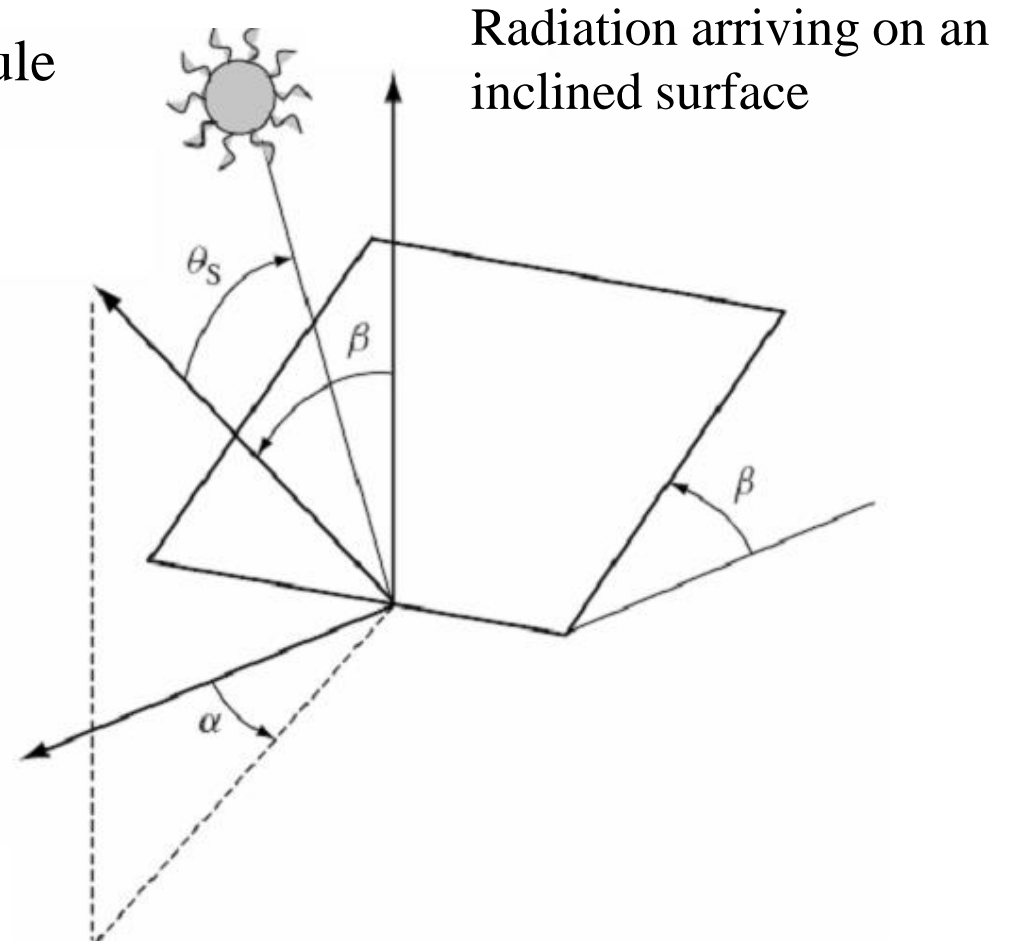
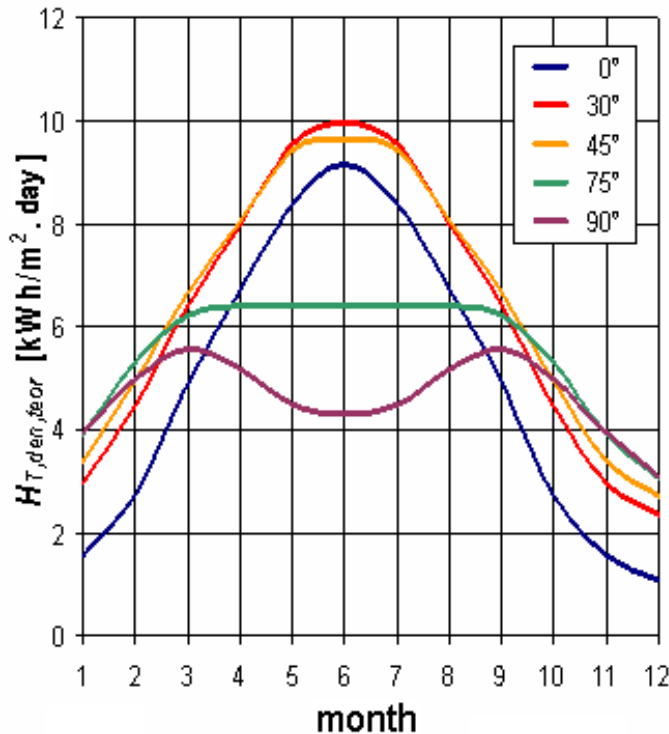


[kWh/(m²a)]





Radiation arriving on PV module

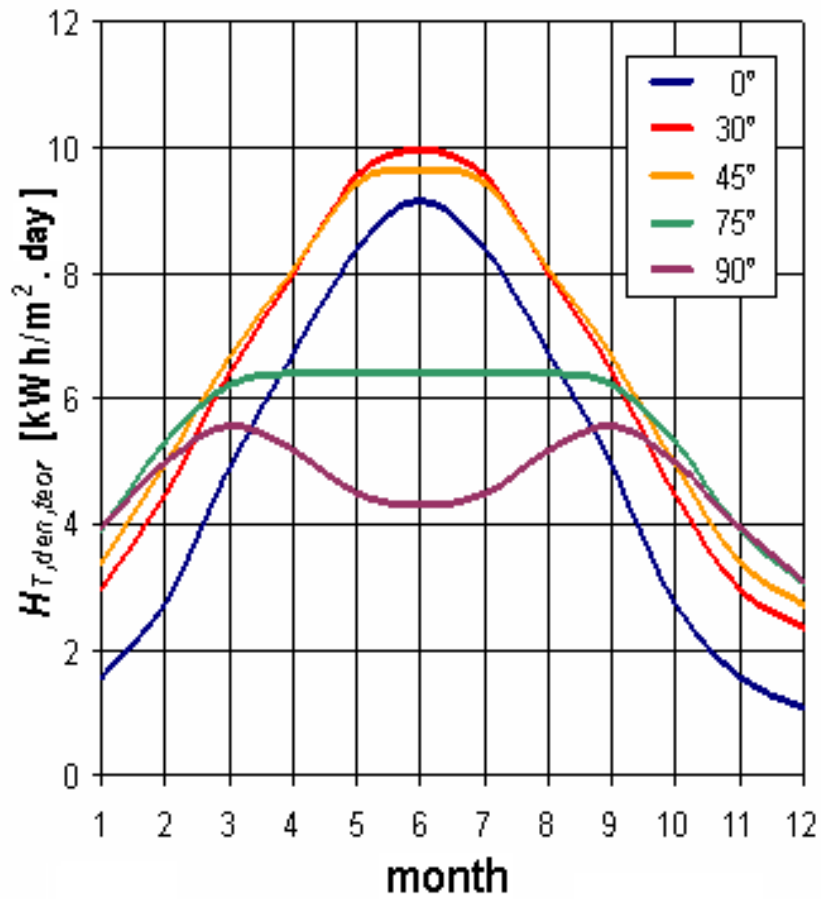


Radiation arriving on an inclined surface

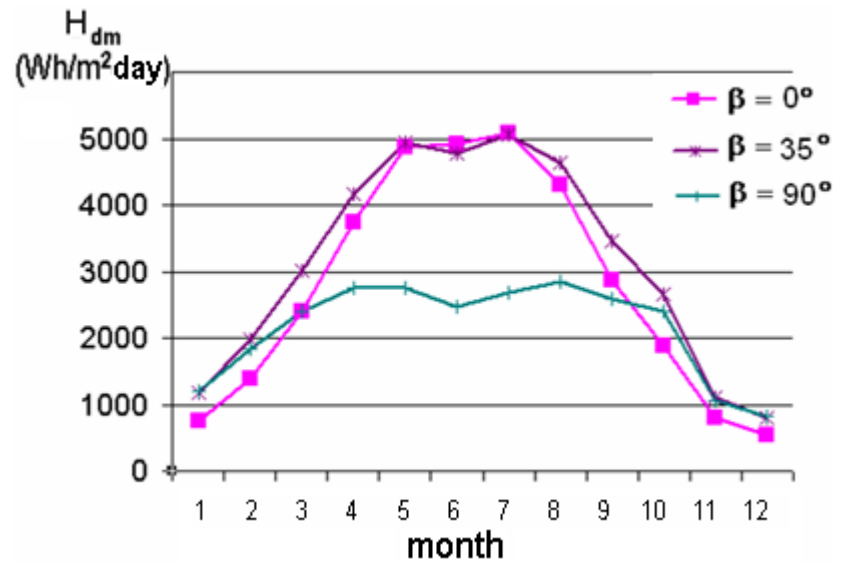
$$\cos \theta_s = \sin \delta \sin \phi \cos \beta - [\text{sign}(\phi)] \sin \delta \cos \phi \sin \beta \cos \alpha + \cos \delta \cos \phi \cos \beta \cos \omega + [\text{sign}(\phi)] \cos \delta \sin \phi \sin \beta \cos \alpha \cos \omega + \cos \delta \sin \alpha \sin \omega \sin \beta$$

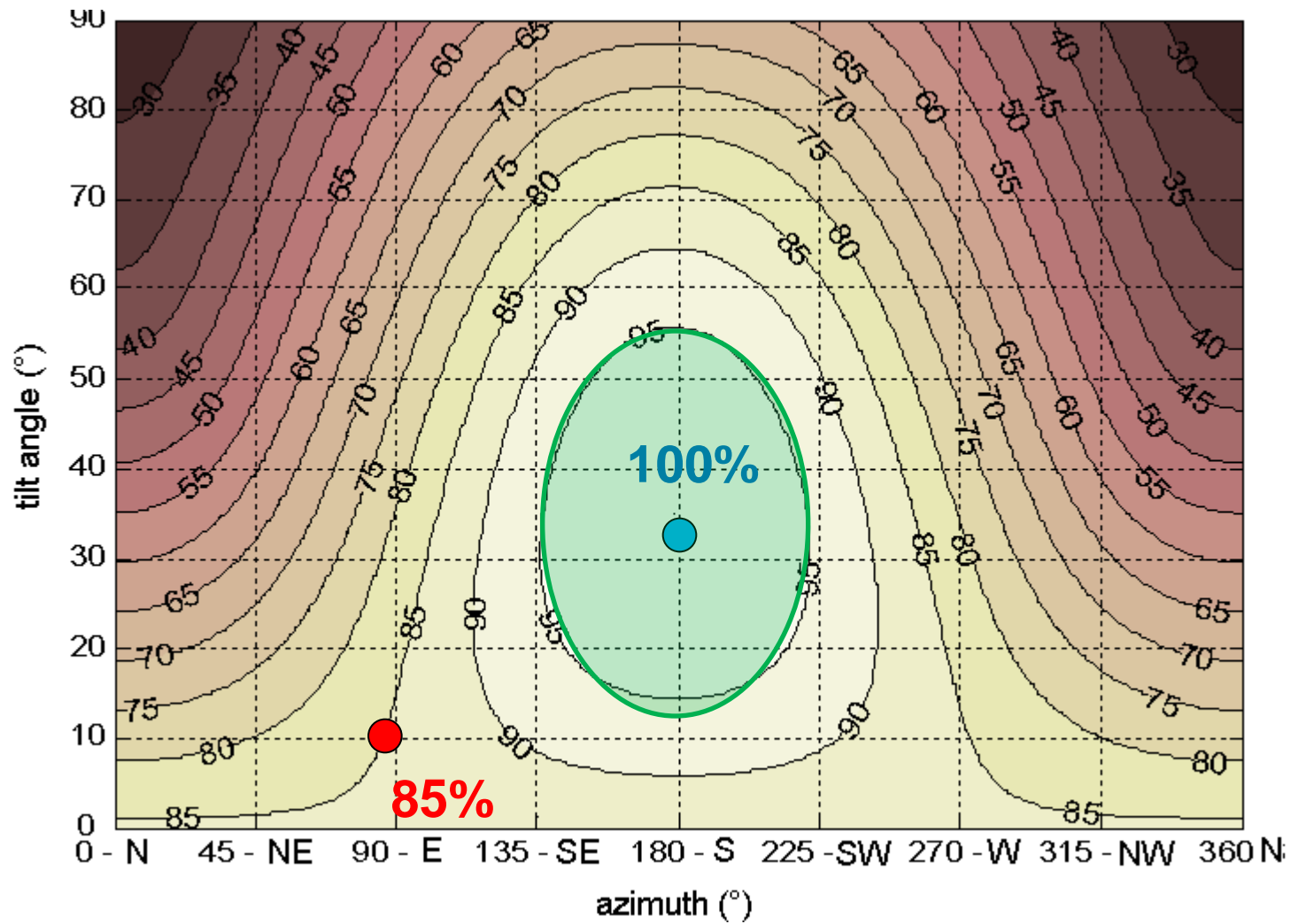
$$\cos \theta_s = [\text{sign}(\phi)] \sin \delta \sin(\text{abs}(\phi) - \beta) + \cos \delta \cos(\text{abs}(\phi) - \beta) \cos \omega$$

Clear sky



Reality:





<http://sunbird.jrc.it/pvgis/apps/pvest.php>

1. Light absorption in materials and excess carrier generation

Absorption is due to interactions with material particles (electrons and nucleus).

If particle energy before interaction was W_1 , after photon absorption is $W_1 + h\nu$

- **interactions with the lattice – low energy photons, results in an increase of temperature**
- **interactions with free electrons - important when the carrier concentration is high, results also in temperature increase**
- **interactions with bonded electrons- the incident light may generate some excess carriers (electron/hole pairs)**